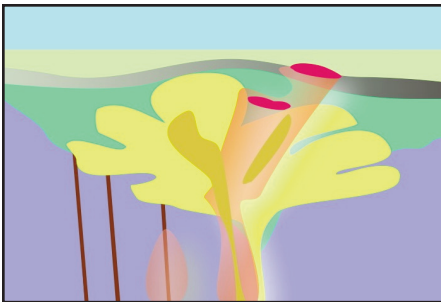


# SERIES



## Igneous Rock Associations 14. The Volcanic Setting of VMS and SMS Deposits: A Review

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### SUMMARY

Volcanogenic massive sulphide (VMS) deposits and seafloor massive sulphide (SMS) deposits have a spatial and genetic connection with contemporaneous volcanism. The control exerted by the volcanic succession (e.g. rock type, architecture and facies) on the nature and style of the ore and alteration (e.g. subsea-floor replacement vs. exhalative, or discordant vs. conforma-

ble) is significant, making it imperative to understand the local volcanology in developing better genetic and exploration models. Three VMS deposit groupings collectively represent a high proportion of cases: (1) deposits associated with complexes of submarine felsic domes, cryptodomes, lobe-hyaloclastite flows and/or blocky lavas, and their reworked equivalents; (2) deposits associated with thick piles of pumiceous felsic pyroclastic rocks, suggesting a caldera context; and (3) deposits associated with mafic volcanic footwalls and/or with sedimentary hosts, including significant deposits such as Windy Craggy (~300 Mt) in British Columbia. With regard to number (2) above, demonstrating the presence of a caldera in ancient successions can be difficult because silicic calderas tend to be large and exceed the limits of deposit-scale investigations. Furthermore, there is no consensus regarding what a large submarine caldera should look like, i.e., no accepted facies model exists showing the distribution of rock types. But without thick piles of pumiceous felsic pyroclastic deposits, arguing for a large submarine caldera is a challenge.

### SOMMAIRE

Les gisements de sulfures massifs volcanogènes (SMV) et leurs équivalents actuels au fonds des mers ont une connexion spatiale et génétique avec le volcanisme. La succession volcanique – composition, architecture, faciès – exerce un contrôle important sur la nature et le style de minéralisation et d'altération hydrothermale (p. ex. minéralisation mise en place par remplacement sous le fond marin vs. exhalative;

altération discordante ou plus concordante). Il est donc impératif de connaître la volcanologie des roches encaissantes pour développer de meilleurs modèles génétiques et d'exploration. Trois groupes de gisements couvrant collectivement une grande proportion des cas sont discutés ici. Premièrement, plusieurs gisements sont associés à des complexes de dômes felsiques sous-marins, des cryptodômes, des coulées de type lobes-hyaloclastite et/ou des laves en blocs, ou leur équivalents resédimentés. Deuxièmement, certains gisements sont associés à d'épaisses séquences de roches pyroclastiques felsiques ponceuses, suggérant un contexte de caldera. Troisièmement, plusieurs gisements sont associés avec des roches volcaniques mafiques et/ou avec des roches sédimentaires, par exemple l'important dépôt de Windy Craggy (~300 Mt) en Colombie-Britannique. Concernant les contextes de type 2, la démonstration d'une caldera peut être difficile dans les successions anciennes, car les caldeiras felsiques sont de grandes dimensions, excédant les limites des études à l'échelle du gîte. De plus, il n'existe pas de consensus sur un modèle de faciès pour une grande caldera sous-marine. Mais sans la présence d'épais empilements de roches pyroclastiques felsiques ponceuses, il est difficile d'argumenter en faveur d'une caldera sous-marine.

### INTRODUCTION

About 300 sites with accumulations of seafloor massive sulphide (SMS) are known on the modern ocean floor (Shanks and Thurston 2012). Of these, 65% are located on mid-ocean ridges,

22% are found in back-arc basins and 12% are within submarine arcs (Hannington et al. 2005). In the future, they may be mined for their base and precious metals. These deposits are the modern equivalent of ancient volcanogenic massive sulphide (VMS) deposits found on land. VMS deposits are “strata-bound accumulations of sulphide minerals that precipitated at or near the seafloor in spatial, temporal and genetic association with contemporary volcanism” (Franklin et al. 2005). They are thought to have been predominantly formed, or at least preferentially preserved, in arc and back-arc settings. VMS deposits represent major sources of Ag, Au, Cu, Pb and Zn (Barrie and Hannington 1999; Allen and Weihed 2002); for example, they account for half of Canada’s Zn production and nearly a third of its Cu (Galley et al. 2007), making them attractive targets for exploration. Historically, many VMS deposits have been found by geophysical and other direct methods of detection; such methods work best at shallow depths. The necessity to explore at greater depth in mature camps forces geologists not only to recognize direct evidence for mineralization, but also favourable host environments. Physical volcanology studies such as those reviewed here can help identify favourable host environments for VMS mineralization. According to Gibson et al. (1999), in the context of VMS exploration, physical volcanology is the study of “(1) the products and deposits formed by volcanic eruptions, (2) eruptive mechanisms, (3) processes of emplacement, and (4) landforms produced by eruptions”.

### Lithostratigraphic and Volcanological VMS Classifications

A number of VMS classification methods were proposed in the past, but the most common current classification uses a “lithostratigraphic scheme based primarily on the principal volcanic and sedimentary lithological units that formed concurrently with the deposits in a given district” (Franklin et al. 2005). This scheme includes five types: (1) bimodal-mafic, (2) mafic, (3) pelitic-mafic, (4) bimodal-felsic, and (5) siliciclastic-felsic (Franklin et al. 2005). Galley et al. (2007) added a sixth group,

‘hybrid bimodal felsic’, for deposits that combine epithermal and VMS characteristics and that are thought to have formed in a shallow water setting.

Building on the work of Mor-ton and Franklin (1987), Gibson et al. (1999) separated VMS deposits hosted by volcanic rocks into two types: ‘flow dominated’ and ‘volcaniclastic dominated’. This field classification is based on the character of the footwall strata, up to 1 km or more below the deposit. According to Gibson et al. (1999), Doyle and Allen (2003), and Franklin et al. (2005), the character of the submarine volcanic rocks, whether coherent lava (massive or pillowed lava flows, lava dome cores) or fragmental (volcaniclastic rocks of all types), influences:

- the size, morphology and growth mechanism of the VMS deposit, i.e., whether it is a mound-shaped sulphide lens formed mostly on the seafloor, with an underlying discordant stockwork, or a tabular sulphide body formed mostly by subsea-floor replacement, with a more localized stringer zone or no stringer zone;
- the size, shape and composition of the proximal footwall alteration, i.e., whether it is a discordant pipe of chlorite-sericite-quartz  $\pm$  carbonate or a broad diffuse zone of sericite, quartz, aluminous silicate and carbonate, with local chlorite; and
- the character of regional semi-conformable alteration zones.

### Importance of Physical Volcanology Studies and Terminology for Volcaniclastic Rocks

The control exerted by the volcanic succession (e.g. rock type, architecture and facies) on the nature and style of the ore and alteration (e.g. subsea-floor replacement vs. exhalative, or discordant vs. conformable) is significant, making it essential to understand the physical volcanology in developing better genetic and exploration models. Physical volcanology studies can help locate synvolcanic structures, which control the location of VMS deposits and can represent coincident effusive and hydrothermal centres (Gibson et al. 1999). In other words, VMS deposits are commonly associated with

proximal volcanic settings (e.g. Gibson et al. 1993; Allen et al. 1996a; Allen and Weihed 2002). Field and drill core studies of lateral and vertical facies variations can help reconstruct the volcanic setting of VMS deposits and identify potentially fertile areas (e.g. McPhie et al. 1993; Gibson et al. 1999; Rogers et al. 2014).

When describing fragmental volcanic rocks we follow the terminology of White and Houghton (2006). In this scheme, there are four types of primary volcaniclastic rocks: pyroclastic (“sedimentation from pyroclastic plumes and currents”), autoclastic (“deposition of fragments from lava, formed via air cooling”), hyaloclastic (“deposition of fragments from lava, formed via water chilling”) and peperitic (“mingling of magma with wet sediment”).

### Scope of This Review

In this paper we review the volcanic settings of VMS and SMS deposits associated with: (1) felsic dome complexes and lavas; (2) pumiceous felsic pyroclastic rocks (submarine calderas); and (3) basaltic volcanic rocks and sills. The first category is probably the most common (Allen and Weihed 2002). However, the first two categories are not mutually exclusive, since felsic domes can occur in calderas, but it is interesting to contrast districts where calderas exist (type 2) with districts where calderas are absent, are being questioned, or have never been proposed (type 1). This contribution is intended as a short review to illustrate the subject, rather than a comprehensive one, therefore many examples and references are omitted.

### VMS DEPOSITS ASSOCIATED WITH FELSIC DOME COMPLEXES AND LAVAS

It has long been known that complexes of submarine felsic domes and/or blocky lavas, as well as cryptodomes, can be spatially associated with VMS deposits, especially of the bimodal-mafic, bimodal-felsic and siliciclastic-felsic types (e.g. Allen 1992; Cas 1992; Doyle and Allen 2003; Franklin et al. 2005). In addition, lobe-hyaloclastite flows, which are more extensive lavas comprising lesser proportions of hyaloclastite (Gibson et al. 1999), can

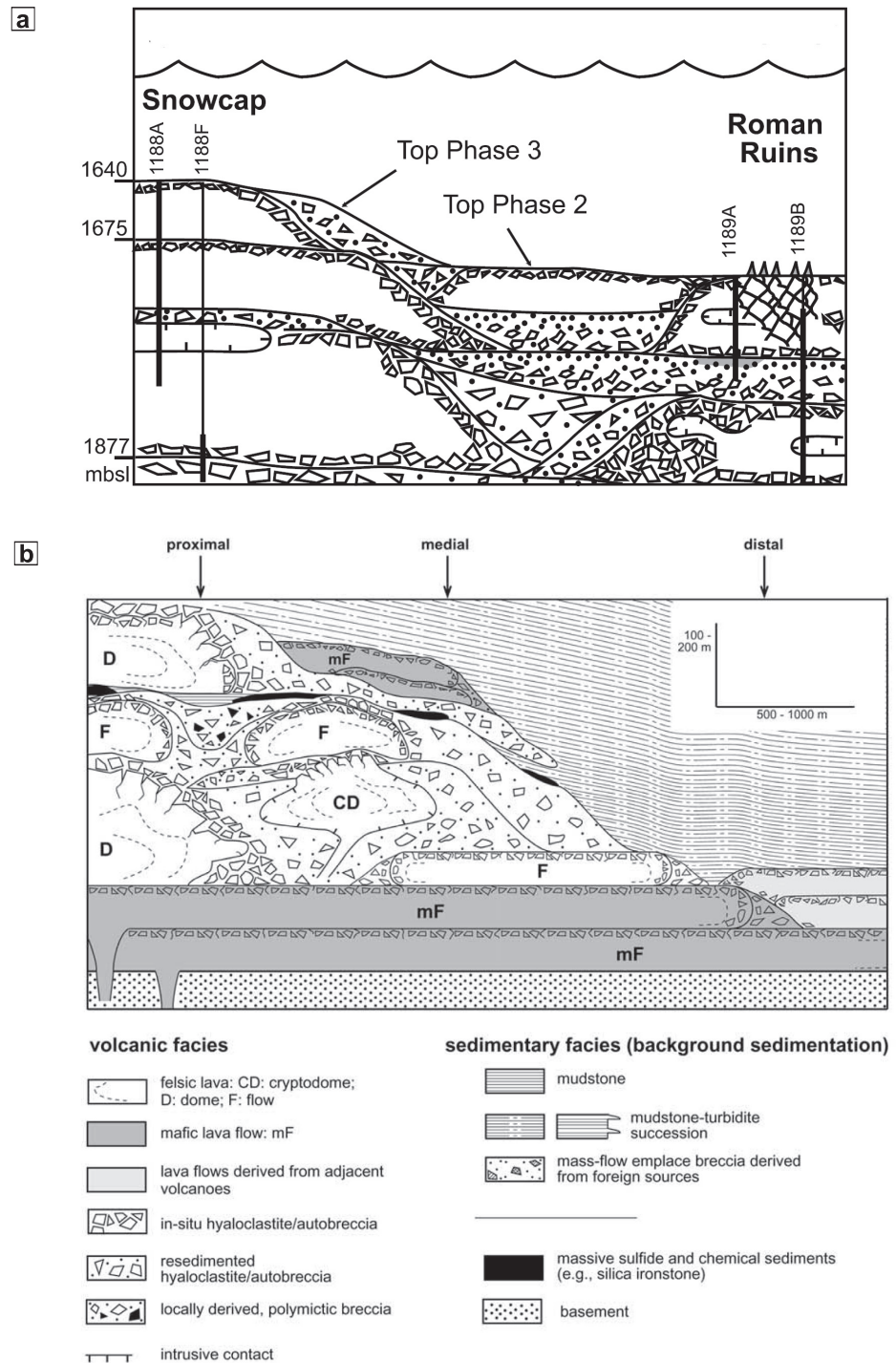
also be spatially related to VMS deposits. Examples from five districts are reviewed, ranging in age from modern to Archean.

**Manus Basin**

The Canadian company Nautilus Minerals has identified 17 SMS prospects in the Manus (back-arc) basin off Papua New Guinea (Lipton 2012). Their ‘Solwara-1’ prospect (a.k.a. Susu Knolls; Binns and Scott 1993) is a candidate for the first underwater mining operation of a SMS deposit, but in terms of physical volcanology, Solwara-4 – formerly called Pacmanus – is better known (Binns and Scott 1993; Scott and Binns 1995). The Ocean Drilling Program obtained several cores of the volcanic rocks and sulphide mineralization from Solwara-4 as part of ODP Leg 193; Paulick et al. (2004) show from these cores that the SMS deposits are directly underlain by a dacitic sequence, which they interpret as a series of 40 – 100 m-thick lava flows and/or domes and associated breccias (Fig. 1). A reconstruction of the volcanic pile shows a number of juxtaposed and superimposed felsic lavas or domes with coherent cores and fragmental margins, as is typical of submarine felsic lavas and domes (e.g. McPhie et al. 1993). Scott and Binns (1995) compared the Solwara-4 site with the Millenbach VMS deposit in the Noranda district (Abitibi greenstone belt), where the ore also sits on top of a complex of felsic domes (see below).

**Hokuroku District**

Several Cu–Zn–Pb–Ag–Au VMS deposits of Miocene age (~15 – 12 Ma; Tanimura et al. 1983; Allen and Weihed 2002; Yamada and Yoshida 2011), locally known as ‘Kuroko’ deposits, have been exploited in the Green Tuff Belt of Japan, including a cluster of 12 deposits in the Hokuroku district of northern Honshu (e.g. Ohmoto 1996). The total past production of Hokuroku district deposits is estimated at ~90 Mt (Yamada, personal communication 2007). The Kuroko deposits developed in a submarine arc undergoing extension (Yamada and Yoshida 2011). Specifically, the Hokuroku district is found in the Kuroko Rift, which contains over 80%



**Figure 1.** (a) Section and (b) facies model showing dacitic lava flows/domes and distribution of massive sulphide deposits at Solwara-4 in the Manus Basin. The two sites drilled, Snowcap and Roman Ruins, are approximately 600 m apart. mbsl = metres below sea level. Figure is from Paulick et al. (2004), with permission from Elsevier (licence 3303690703269).

felsic rocks (Yamada and Yoshida 2011), and the VMS deposits are of the bimodal-felsic type (Franklin et al. 2005).

It has been suggested that a number of submarine calderas occur in

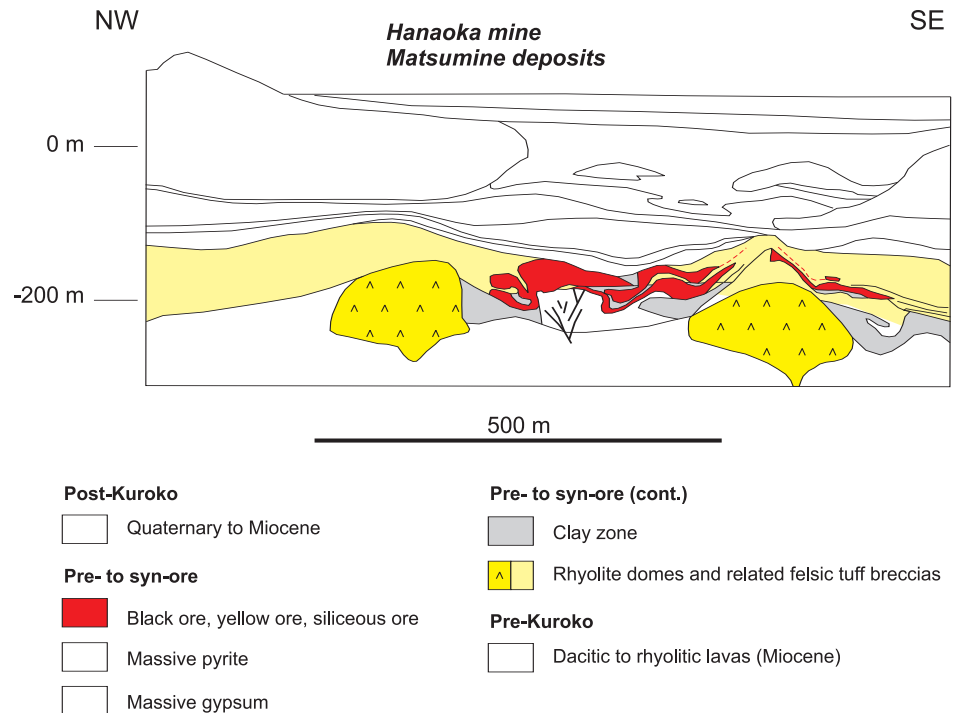
the Hokuroku district (e.g. Ohmoto and Takahashi 1983; Ohmoto 1996; and references therein) and that felsic pyroclastic rocks are abundant in the district (e.g. Ohmoto and Takahashi 1983; Tanimura et al. 1983). However,

Cas (1992) proposed that fragmental rocks around the Kuroko deposits are predominantly hyaloclastite, rather than pyroclastic rocks. He also stated that typical caldera-filling deposits (submarine ignimbrites) are absent in the Hokuroku district and that caldera-bounding faults have not been recognized. However, dacitic to rhyolitic domes and associated breccias are closely associated with the Kuroko ores (Horikoshi 1969; Cas 1992). For example, Kuroda (1983) shows a model for the Furutobe deposit in which the ore is formed in volcanoclastic rocks near a dome. In Yamada and Yoshida's (2011) cross-section of the ~30 Mt Matsumine deposit, the ore lies between two rhyolite domes (Fig. 2). It seems clear that the VMS deposits of the Hokuroku district are associated with lava domes, and not with submarine calderas (Cas 1992; Allen and Weihed 2002).

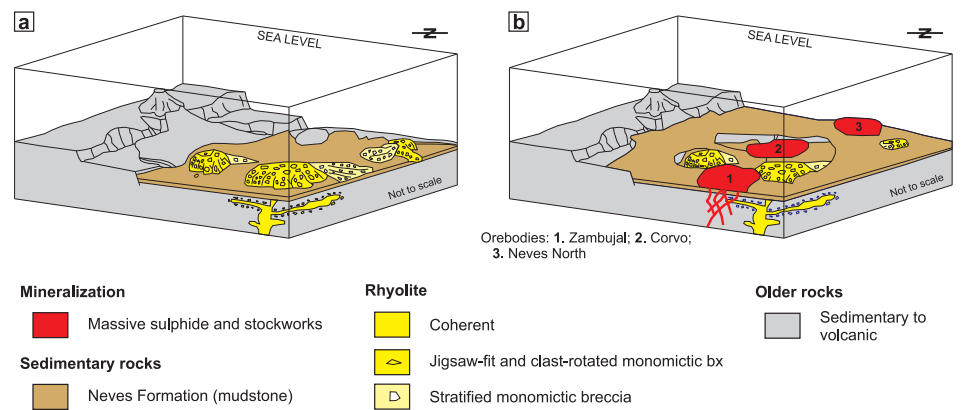
### Iberian Pyrite Belt

About 90 VMS deposits occur in the Iberian pyrite belt of Spain and Portugal, of which seven or eight are larger than 100 Mt, making it one of the most important VMS regions of the world (Allen and Weihed 2002; Rosa et al. 2010). Mercier-Langevin et al. (2014) calculate that the Iberian pyrite belt contains 23% of the global VMS tonnage. The rocks hosting the VMS deposits are Devonian to Carboniferous in age; the succession is dominated by sedimentary rocks, especially mudstones, but mafic to felsic volcanic rocks are also present (Tornos 2006). Soriano and Marti (1999) concluded that volcanism in the eastern part of the Iberian pyrite belt was mostly non-explosive. Rosa et al. (2010) reviewed felsic volcanic centres in the belt and concluded that "these volcanoes are dominated by felsic lavas/domes that occur at several stratigraphic positions... however the pyroclastic units are also abundant".

Rosa et al. (2008) made a detailed study of the volcanic rocks near the 300 Mt (Relvas et al. 2006) Neves Corvo deposit in Portugal and convincingly showed that the immediate footwall includes a coherent to fragmental rhyolite unit forming a series of lava domes or flows (Fig. 3). The total thickness of rhyolite is typi-



**Figure 2.** Vertical cross-section at the Hanaoka mine, Hokuroku district, Japan, showing the association between lava domes and Kuroko deposits. Simplified from Yamada and Yoshida (2011).



**Figure 3.** Schematic diagrams of the Neves Corvo area showing: (a) emplacement of rhyolite lavas/domes on a mudstone substrate; (b) deposition of VMS ore bodies and more mudstone. Modified from Rosa et al. (2008).

cally 85–135 m. The rhyolitic rocks are highly altered and have the same age as the ores (Rosa et al. 2008). Relvas et al. (2006) show photos of sulphides replacing rhyolitic rocks, probably of a fragmental nature. The domes and flows were emplaced in a submarine epicontinental sedimentary basin (Rosa et al. 2008).

### Skellefte District

The Skellefte district of northern Sweden is of Paleoproterozoic (1.90–1.87 Ga) age and contains about 85 VMS

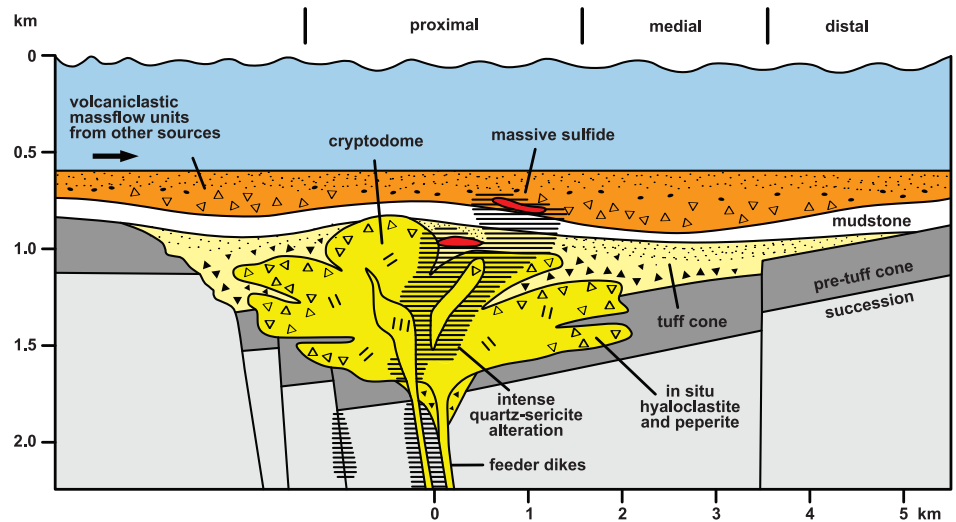
deposits and occurrences (Allen et al. 1996a), of which 31 have been put into production (Mercier-Langevin et al. 2013). In the 1980s, the felsic volcanic rocks in the district were thought of as pyroclastic and related to large calderas, but Allen et al. (1996a) note that "many volcanic rocks in the district, including many originally homogeneous intrusions and lavas, have streaky, patchy, pseudotuffaceous textures" due to diagenetic compaction and tectonic fabrics. Instead of being found in calderas, several VMS

deposits are closely associated with a specific type of volcano called “subaqueous rhyolite cryptodome-tuff volcano” (Allen et al. 1996a). This type of volcano occupies 5% or less of the Skellefte district but many VMS deposits occur in such rocks, or above them (Fig. 4). According to Allen et al. (1996a), “these rhyolite volcanoes are 2 to 10 km in diameter, 250 to 1200 m thick at the center, and are characterized by a small to moderate volume rhyolitic pyroclastic unit, intruded by rhyolite cryptodomes, sills, and dikes. Massive sulphide ores occur near the top of the proximal (near vent) facies association”. The first magma batch was gas-rich and erupted explosively to form the pyroclastic unit (the ‘tuff cone’ in Fig. 4), which is thickest in the vent area (50–300 m). The pyroclastic unit is thickly bedded, pumiceous, and probably accumulated in depressions on the seafloor. It was then intruded by gas-poor magma which formed “sills, dikes, cryptodomes, partly emergent cryptodomes, and possibly some lavas” (Allen et al. 1996a).

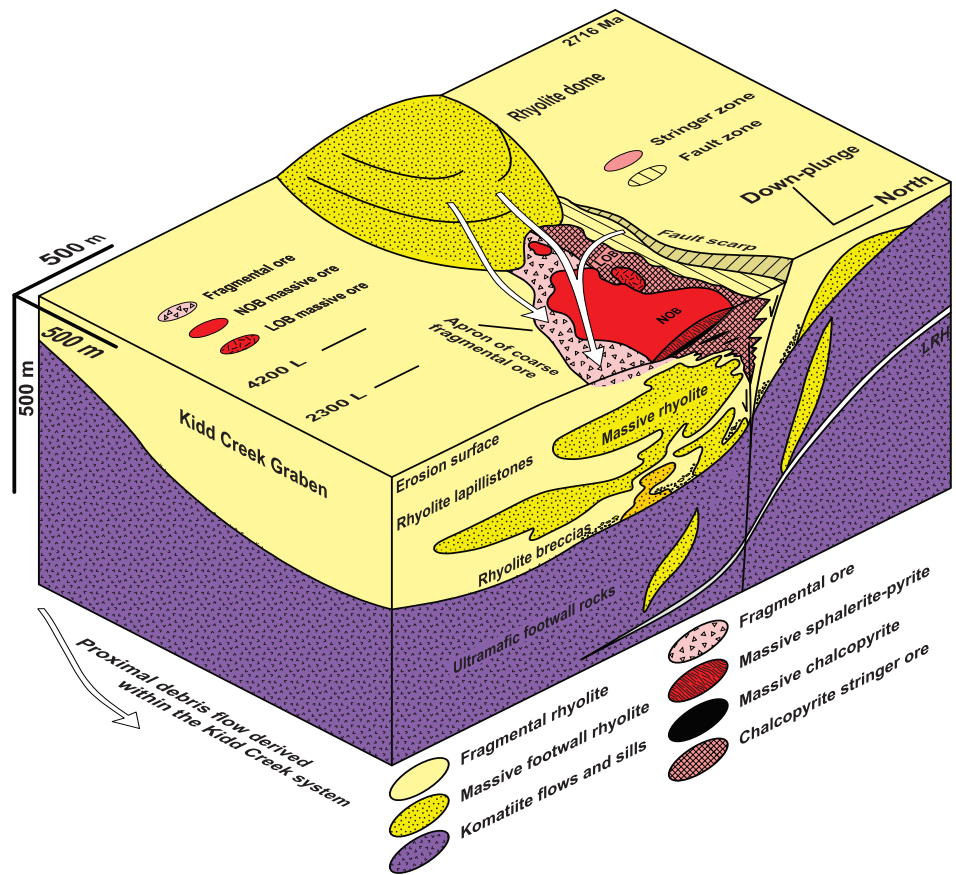
**Abitibi Greenstone Belt**

The Archean Abitibi greenstone belt of Quebec and Ontario contains a VMS tonnage of over 800 Mt (Galley et al. 2007; Goutier et al. 2011; Mercier-Langevin et al. 2011). We focus on two VMS areas from the southern Abitibi belt: Kidd Creek (a large isolated deposit) and the Noranda district. A mention is also made of recent work in the Matagami district.

The Kidd Creek deposit in Ontario is classified as bimodal-mafic (Galley et al. 2007) with ~185 Mt of ore grading 2.3% Cu, 6.2% Zn, 0.2% Pb, 74 g/t Ag, and 0.01 g/t Au (Goutier and Bécu, unpublished compilation, 2013). The deep footwall consists of komatiitic rocks, whereas the immediate footwall is a ~300 m-thick rhyolitic sequence (Bleeker 1999; Hannington et al. 1999). The lower part of the rhyolite sequence is mostly coherent and interpreted as domes and lavas, whereas the upper part is dominated by bedded volcanoclastic rocks inferred to fill a graben (Bleeker 1999). These volcanoclastic rocks are interpreted as mostly derived from felsic domes or lavas as well (Fig. 5). The ores formed mostly by subsea-floor replacement of these



**Figure 4.** Facies model for ‘subaqueous rhyolite cryptodome-tuff volcanoes’ in the Skellefte district. Modified from Allen et al. (1996a).



**Figure 5.** Block diagram showing the initial development of the Kidd Creek VMS deposit. A rhyolite flow-dome complex is emplaced in a subsiding graben or half-graben. The structure is then filled by volcanoclastic debris from sources including the flow-dome complex. Hydrothermal fluids permeate this fragmental pile and VMS orebodies form below the seafloor. Figure modified from Bleeker (1999). Note the x2 vertical exaggeration. LOB = Lowermost orebody; NOB = North orebody.

fragmental rocks (Bleeker 1999; Hannington et al. 1999).

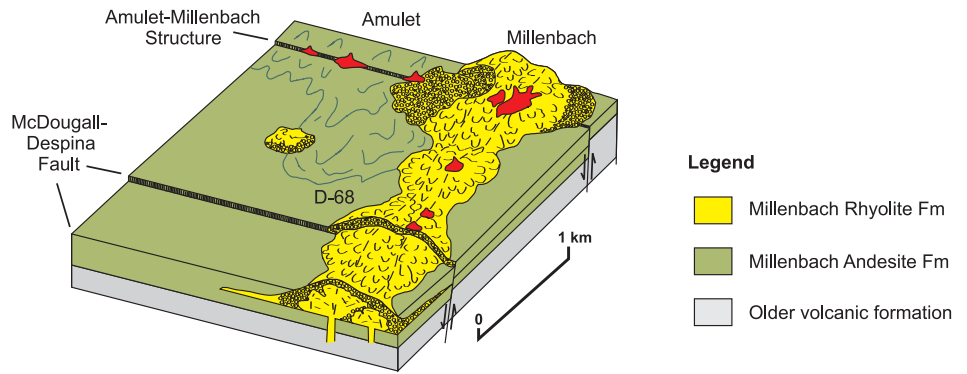
In the central part of the Noranda district (Quebec), 17 small

(typically ~0.5–4.5 Mt) bimodal-mafic, Cu–Zn deposits occur in a lava-dominated volcanic succession (Gibson and Watkinson 1990). The Rouyn-Noranda area has abundant outcrop and is a classic locality for submarine volcanology and VMS studies. Two types of rhyolitic volcanoes are known in the Noranda district, blocky lavas/domes and lobe-hyaloclastite flows (de Rosen-Spence et al. 1980; Gibson et al. 1999) and both can be associated with VMS mineralization. The Millenbach mine area is a good example of VMS mineralization associated with a blocky rhyolitic lava/dome. This dome/flow is described as a “1.8 km-long northeast-trending ridge that directly overlies its feeding fissure” (Gibson et al. 1999). The maximum thickness of rhyolite is 250 m and the slopes of the ridge range from 30 to 70°. The interior of the ridge consists of coherent rhyolite whereas the top and sides consist of breccias (Fig. 6). The breccias are up to 30 m thick (Gibson et al. 1999). Ore occurred “along the top of the rhyolite ridge and directly above and along the feeding fissure” (Gibson and Galley 2007), forming a series of 15 VMS bodies up to ~1 Mt in size (Gibson et al. 1999). Knuckey et al. (1982) show a ‘QFP dome’ isopach map supporting this interpretation.

In contrast, ore occurred immediately above lobe-hyaloclastite flows at the Ansil, Vauze, and Norbec deposits in the Noranda district (Gibson et al. 1999). Lobe-hyaloclastite flows comprise a lower proportion of fragmental rocks relative to blocky domes/flows. They form “broad, gentle-sloped (10°–20°) lava shields or plateaus that attained heights of up to 500 m”, are fed by fissures, and can reach 5 km in radius (Gibson et al. 1999). Recent studies in the Matagami district, in the northern part of the Abitibi greenstone belt, suggest that lobe-hyaloclastite flows are also the dominant type of felsic lavas immediately below and above the Zn-rich VMS deposits of the district (Debreil 2014).

### VMS DEPOSITS ASSOCIATED WITH PUMICEOUS FELSIC PYROCLASTIC ROCKS

Some SMS and VMS deposits are associated with pumiceous felsic pyroclastic



**Figure 6.** Reconstruction of the ‘Millenbach dome complex’, which is a rhyolite ridge found above its feeding fissure, in the Noranda district. This ridge consists of a series of blocky domes/lava flows, upon which is a series of VMS ore lenses (red). Modified from Gibson and Galley (2007).

rocks that formed in a submarine caldera setting. Brothers volcano, in the Kermadec-Tonga arc, north of New Zealand, is an example of a submarine caldera that has SMS mineralization forming in it today (de Ronde et al. 2005). Several scientific cruises have visited this volcano and an active hydrothermal field has been mapped (Baker et al. 2012 and references therein). Ancient examples of VMS deposits in felsic submarine calderas are found in the Proterozoic Bergslagen district of Sweden (Allen et al. 1996b) and the Archean Sturgeon Lake district of Ontario (bimodal-felsic type; Hudak et al. 2003 and references therein).

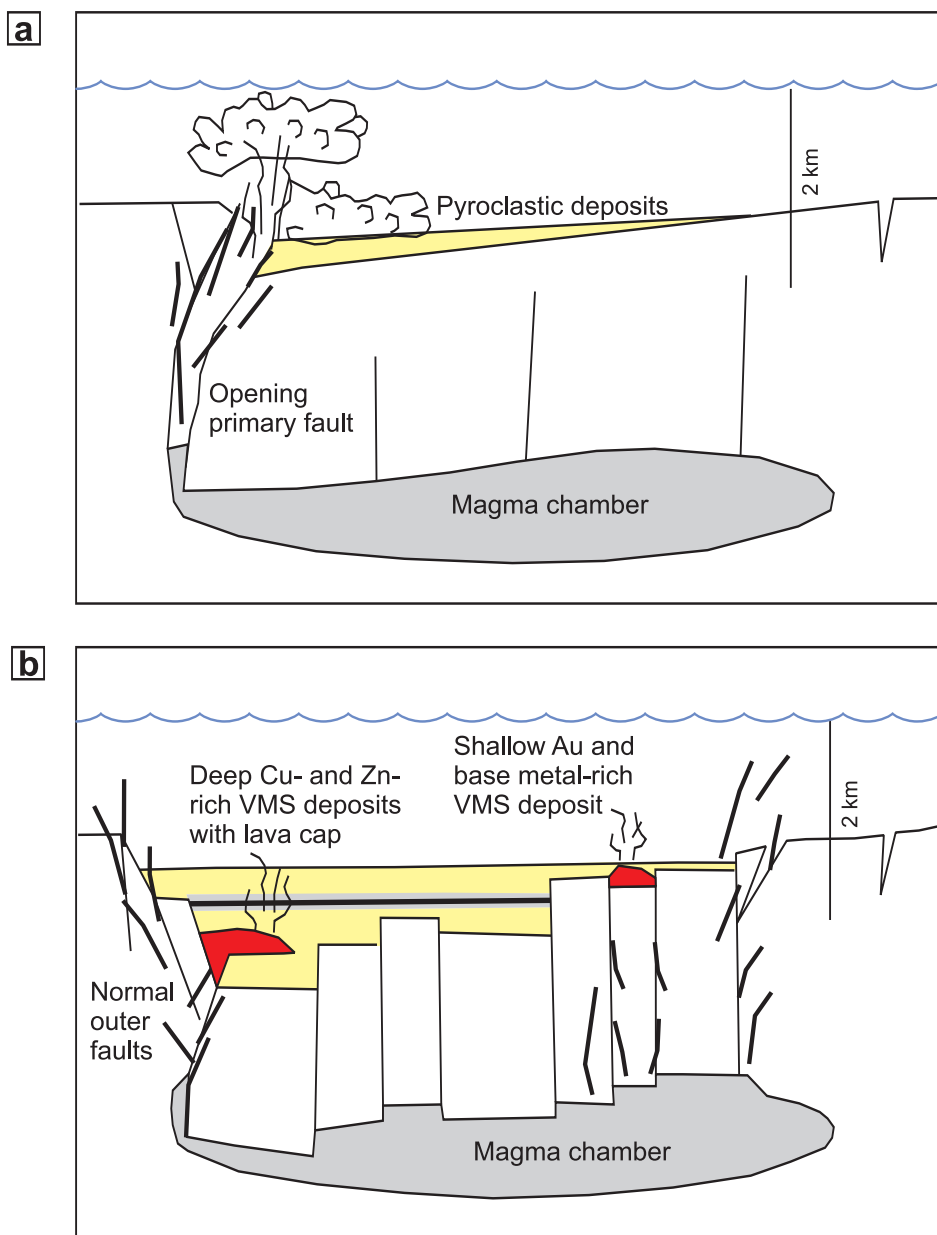
### Submarine Calderas

Because submarine calderas are not well known, we start by describing their continental equivalents. The largest continental calderas – which can reach tens of kilometres in diameter – are associated with catastrophic explosive eruptions of gas-rich felsic magmas (e.g. Cole et al. 2005 and references therein). The top of the magma chamber empties rapidly because of the catastrophic explosive eruption, and this leads to caldera collapse (Smith and Bailey 1968). The typical products of such eruptions are pumiceous pyroclastic flow deposits (ignimbrites). Within the caldera, these deposits can be up to kilometres in thickness, but outside the caldera they are much thinner (Lipman 1997). Very thick sequences of pumiceous, poorly-sorted felsic pyroclastic rocks are thus compatible with a caldera setting, even if the original caldera geometry cannot

be recognized in ancient rocks.

Submarine calderas are less known, but the largest ones also seem to be associated with large explosive eruptions of felsic magma (e.g. Fiske et al. 2001; Yuasa and Kano 2003) and form by the same mechanism of roof collapse above a rapidly emptying magma chamber. The felsic calderas of the northern Isu-Bonin arc range in diameter from 2 to 10 km, which is relatively small compared to continental calderas, and the explosive eruptions create large volumes of pumice (Yuasa and Kano 2003). These felsic submarine calderas occupy a large area of the inferred original volcanic edifice, unlike mafic calderas, in which the collapse area represents only a small portion of the volcanic edifice (e.g. subaerial shield volcanoes in Hawaii). Mafic calderas form as a result of magma withdrawal (e.g. Walker 1993).

Stix et al. (2003, and references therein) proposed that submarine calderas should be excellent hosts for VMS deposits, especially near the caldera margin, because “caldera opening along outward-dipping faults promotes magma degassing, seawater influx, and high-temperature leaching, resulting in a metal-rich hydrothermal fluid” (Fig. 7). Furthermore, “the accumulation of significant thicknesses of pyroclastic deposits within the caldera is an important source of stratigraphic permeability, which can be exploited by mineralizing solutions” to form VMS deposits by subsea-floor replacement (Stix et al. 2003). Two good examples, one modern from the Isu-Bonin arc, and one ancient from northern Maine,



**Figure 7.** Sketch showing the development of a submarine caldera and associated VMS deposits (red): (a) influx of seawater into the magma chamber; (b) development of hydrothermal cells. Modified from Stix et al. (2003).

are now presented in some detail.

**Isu-Bonin Arc**

The Myojin Knoll caldera in the Isu-Bonin arc south of Japan contains an actively forming SMS deposit that was discovered in the 1990s (Iizasa et al. 1999). The 6 x 7 km Myojin Knoll caldera occupies the centre of a 950 m-high, 19 x 22 km diameter submarine felsic volcano (Fiske et al. 2001). A 2 km-diameter, 250 m-high, post-caldera felsic dome occupies the centre of the caldera floor. The collapse volume is ~18 km<sup>3</sup> and the caldera is sur-

rounded by a non-stratified to poorly-stratified pumice layer, up to 200 m thick, which thins outward from the caldera margin (Fig. 8). Fiske et al. (2001) propose that the caldera formed as a result of a catastrophic felsic explosive eruption. Following caldera collapse, massive sulphides began accumulating in the eastern part of the caldera floor, along the inferred trace of the main caldera fault. The Sunrise deposit (Fig. 8) is 400 x 400 m in map view, and is reportedly 30 m thick, generating an estimate of 9 Mt of massive sulphides that are exceptionally rich in

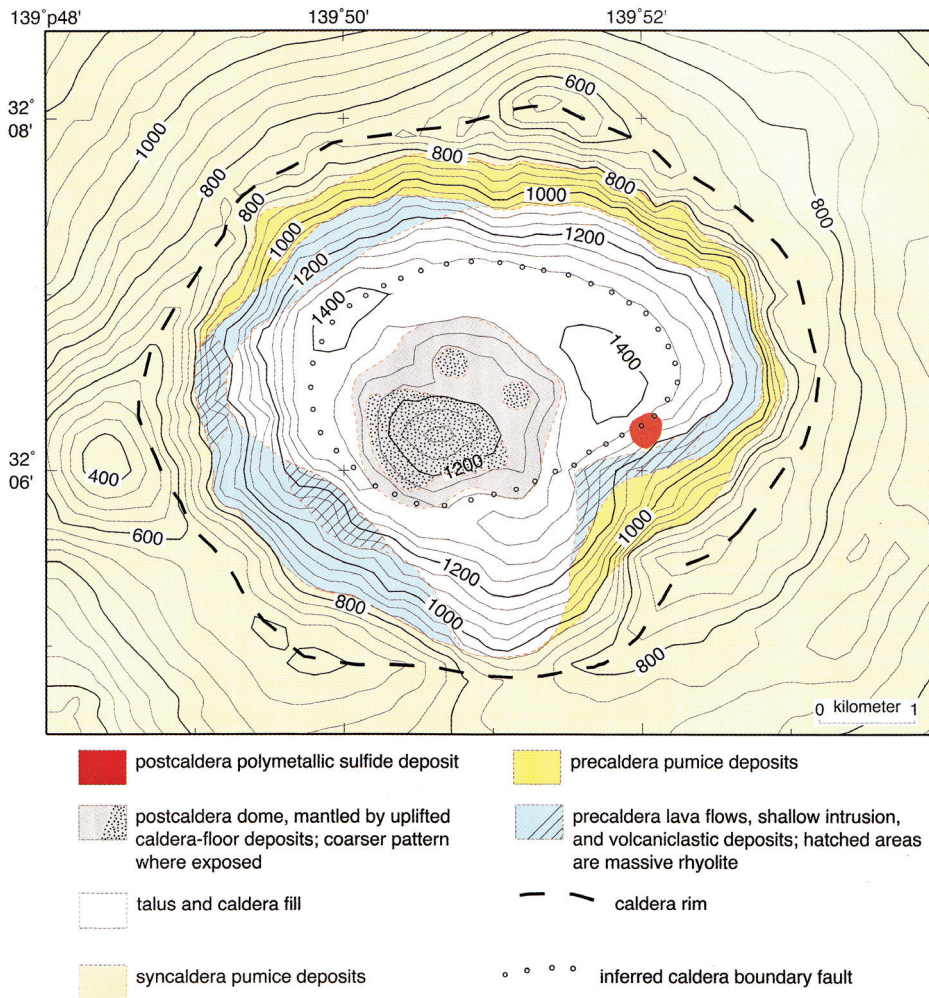
base and precious metals (Iizasa et al. 1999), although the deposit has not been drilled to confirm this thickness and tonnage.

**Bald Mountain**

The Ordovician Bald Mountain VMS deposit in northern Maine, a bimodal-felsic type, contains 30 Mt of hypogene ore at average grades of 1.0% Cu, 1.1% Zn, 0.4% As, 0.5 g/t Au, and 14 g/t Ag (Slack et al. 2003). The metamorphic grade is sub-greenschist, cleavage is lacking, and primary textures are well preserved (Busby et al. 2003; Slack et al. 2003), making this area ideal to study the volcanic host rocks of a VMS deposit. The inferred setting for the volcanic and hydrothermal activity is a deep water (>1.5 km) graben in a primitive volcanic arc (Busby et al. 2003). The deep footwall consists of at least 2 km of submarine tholeiitic basalts and basaltic andesites (lavas) (Fig. 9a). The immediate footwall and hangingwall is a thick ignimbrite succession that contains the VMS deposit. The ~200 m-thick footwall ignimbrite, a crystal-poor, non-welded, non-stratified, pumiceous felsic lapillistone, formed as a result of a major underwater explosive eruption that formed the Bald Mountain sub-basin (Busby et al. 2003), which is possibly a caldera (Fig. 9a). Subsequently, the VMS deposit (up to 215 m thick) formed in the inferred vent of the crystal-poor ignimbrite (Busby et al. 2003); sulphides precipitated through a combination of chimney growth and subsea-floor replacement (Slack et al. 2003). Finally, the ≤350-m thick crystal-rich hangingwall ignimbrite sequence (including andesitic lava intercalations) formed and covered the Bald Mountain deposit (Fig. 9b). This mostly explosive eruption was accompanied by the collapse of the Bull Hill sub-basin, another possible caldera (Busby et al. 2003).

**VMS DEPOSITS ASSOCIATED WITH BASALTS**

Ophiolite-hosted VMS deposits in Cyprus and Oman are enclosed in basaltic volcanic rocks (Hannington et al. 1998). Another good example is the Jurassic Turner-Albright VMS deposit in Oregon (~3 Mt), where the bulk of the mineralization formed by replace-



**Figure 8.** Geological map of the Myojin Knoll caldera showing the distribution of syn-caldera pumice and the location of the Sunrise SMS deposit (“postcaldera polymetallic sulphide deposit”). Republished from Fiske et al. (2001) with permission of the Geological Society of America; permission conveyed through Copyright Clearance Center Inc., licence 3303720448684.

ment of basaltic hyaloclastite (Zierenberg et al. 1998). On the modern seafloor, SMS deposits such as TAG are actively forming in a mafic setting (Hannington et al. 1998).

In Canada, the Windy Craggy VMS deposit of northern British Columbia (pelitic-mafic type), is an excellent example of a very large VMS deposit (~300 Mt) within a succession of Triassic sedimentary and mafic volcanic to intrusive rocks (Fig. 10; Peter and Scott 1999). The host rocks consist of “interbedded pillow basalt and graphitic to calcareous siltstone and argillite” with local basaltic sills (Peter and Scott 1999).

## DISCUSSION AND CONCLUSIONS

Three categories of SMS and VMS

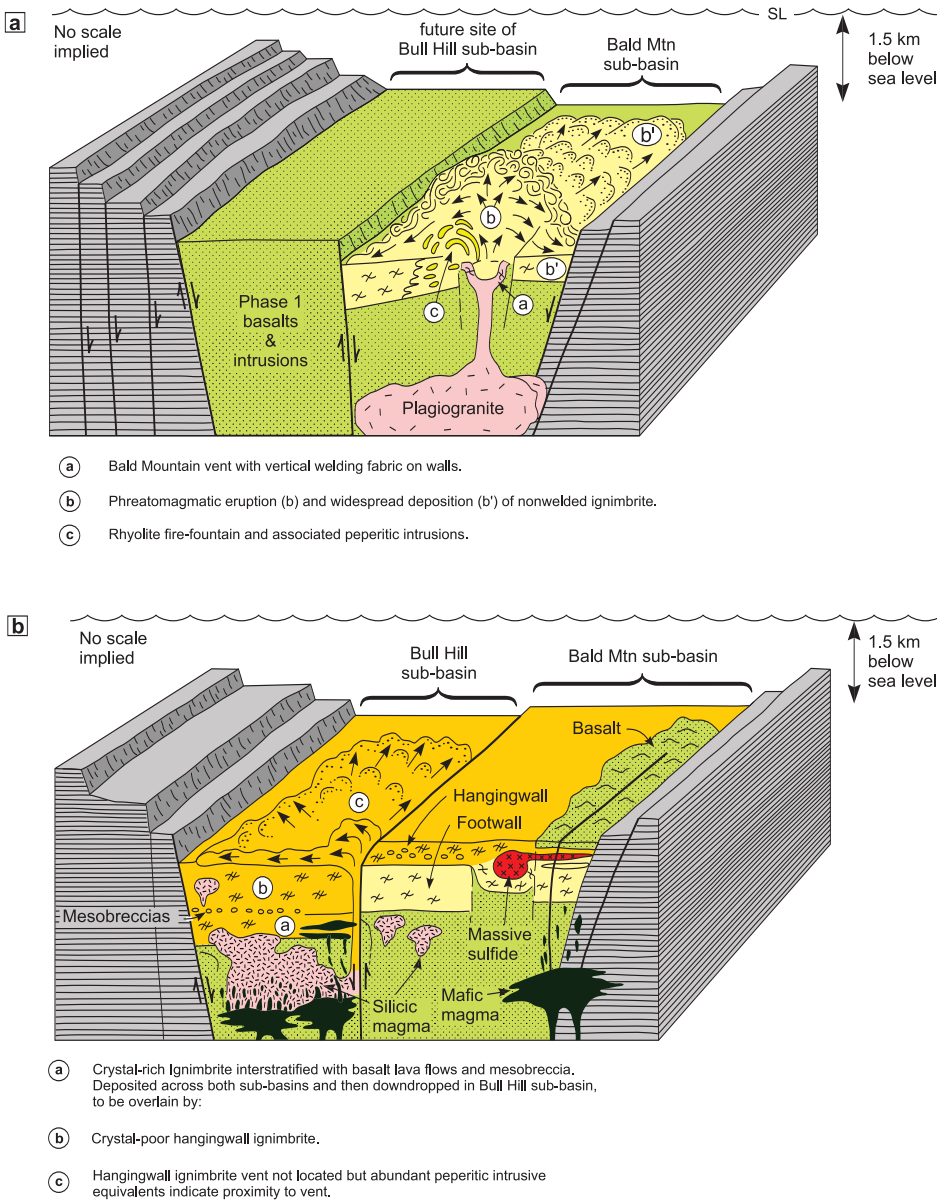
deposits are described in this review: (1) deposits associated with felsic dome complexes and lavas; (2) deposits associated with pumiceous felsic pyroclastic rocks in submarine calderas; (3) deposits associated with basaltic volcanic rocks. However, some deposits are associated with other volcanic settings. For example, Cas (1992) cites examples of Australian VMS deposits that are found in successions dominated by felsic “mass-flow deposits of resedimented pyroclastic debris” (see also McPhie and Allen 1992). Doyle and Allen (2003) give examples of other types of volcanoclastic deposits as well. Nevertheless, the three categories or groups presented here cover a large proportion of cases.

For this review, it was easier to

find detailed and convincing volcanological studies of group 1 deposits than of group 2 deposits. One reason may be that the facies architecture of felsic domes and lavas is well known, relatively simple, and distinctive (e.g. McPhie et al. 1993). A felsic dome is limited in size and the spatial association with a VMS deposit should therefore be clear on local geological maps and cross-sections. Providing primary textures and structures are sufficiently preserved, a VMS deposit that is associated with a felsic dome or lava should be recognizable without much debate.

In contrast, demonstrating a caldera association in ancient successions can be difficult. First, there is a scale challenge: silicic calderas tend to be large, so a regional study is necessary to identify one. Good exposure or extensive drilling coverage are therefore essential, but are not always available. Structural complexities can hamper the recognition of potential calderas in ancient successions. Several authors have emphasized great thicknesses (100s of m to several km) of ignimbrite in the caldera, but distinguishing submarine pyroclastic flow deposits from other types of subaqueous pumiceous deposits is not straightforward and has led to much debate in the literature. One reason is that welded submarine pyroclastic rocks are very rare (Cas 1992; White 2000), so there are few obvious diagnostic criteria to distinguish primary from resedimented deposits. Glassy pumiceous deposits are especially prone to modification by diagenesis, hydrothermal alteration, tectonic deformation, and metamorphism (e.g. Gifkins et al. 2005). Also, a wide range of other volcanoclastic and even coherent volcanic rocks can acquire false pyroclastic textures (Allen 1988), leading to misidentification. Finally, there is no consensus regarding what a submarine caldera should look like, i.e., there is no accepted facies model showing the distribution of rock types.

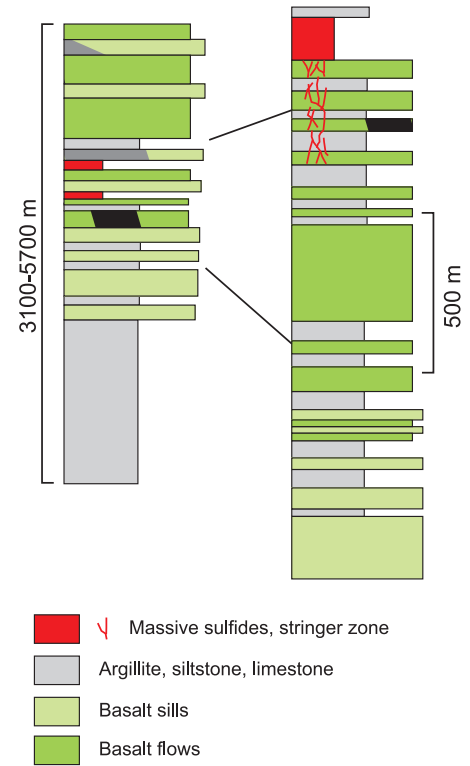
Consequently, while there are certainly good examples of VMS or SMS deposits associated with calderas, such as the two reviewed above, other reported calderas in Australia and Japan (e.g. Ohmoto and Takahashi 1983) have been disproved or doubted in later studies (see Cas 1992; Allen



**Figure 9.** Schematic block diagrams showing the evolution of the Bald Mountain sequence: (a) eruption of footwall ignimbrite and collapse of the Bald Mountain sub-basin, followed by growth of the VMS deposit within the ignimbrite vent; (b) collapse of the Bull Hill sub-basin and eruption of the hanging wall ignimbrite, covering the VMS deposit. Modified from Busby et al. (2003).

and Weihed 2002). Yet following the discovery of the Sunrise SMS deposit in the late 1990s, the caldera model regained strength and classic VMS districts were reinterpreted as calderas or caldera complexes. For example, Stix et al. (2003) cite the Horne VMS deposit (Noranda district) as an example of a deposit located in a caldera margin environment, following the ideas of Gibson and Watkinson (1990) and Kerr and Gibson (1993). However, McNicoll et al. (2014) show through high-precision U–Pb geochronology

that the ~54 Mt Horne deposit (plus ~170 Mt of subeconomic sulphide in Zone 5; Mercier-Langevin et al. 2011) and the neighbouring ~14 Mt Quemont deposit lie in rocks that are about 4 m.y. older than those hosting the ‘central camp’ VMS deposits (such as Millenbach and Amulet) in the inferred centre of the Noranda cauldron, farther north. Therefore, Horne and Quemont do not seem to be related to a caldera margin setting. Taking it a step further, Mueller et al. (2009) proposed that the Bouchard-Hébert, Horne, and



**Figure 10.** Schematic stratigraphic section through the Windy Craggy deposit. Modified from Peter and Scott (1999) and Franklin et al. (2005).

Quemont VMS deposits were associated with nested calderas in the Blake River Group of the Abitibi greenstone belt and that the calderas were comparable to modern “submarine arc calderas such as Myojin Knoll”. However, Mueller et al. (2009) also proposed that these Archean examples were “controlled predominantly by effusive-dominated caldera subsidence rather than explosive magma evacuation (e.g. Myojin Knoll caldera)” so it is not clear what they really meant.

Without thick piles of pumiceous felsic pyroclastic deposits – which ought to be at least partly preserved in a subaqueous environment – arguing for a large submarine caldera (e.g. Pearson and Daigneault 2009; Mueller et al. 2012) is difficult (see also Ross et al. 2011a, b for more discussion on the Blake River Group). Recall that in subaerial settings, calderas produced by magma withdrawal or lava effusions are summit calderas many times smaller than the shield volcanoes or stratovolcanoes on which they occur. In summary, demonstrating the association between submarine calderas

and VMS deposits in ancient successions is difficult, which may be why there are only a few truly convincing examples; alternatively, this association may be rarer than that between VMS deposits and felsic lava domes and flows.

Finally, we stress again that mafic volcanic rocks and sedimentary rocks can also be excellent hosts for VMS deposits (e.g. ~300 Mt at Windy Craggy).

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